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SPIN 1/2 PARTICLE IN THE FIELD OF THE DIRAC STRING ON THE BACKGROUND OF DE SITTER SPACE–TIME

The Dirac monopole string is specified for de Sitter cosmological model. Dirac equation for spin 1/2 particle in presence of this monopole has been examined on the background of de Sitter space-time in static coordinates. Instead of spinor monopole harmonics, the technique of Wigner D -functions is used. After separation of the variables, detailed analysis of the radial equations is performed; four types of solutions, singular, regular, in- and out- running waves, are constructed in terms of hypergeometric functions. The complete set of spinor wave solutions $\Psi_{\varepsilon,j,m,\lambda}(t, r, \theta, \phi)$ has been constructed, special attention is given to treating the states of minimal values of the total angular momentum j_{\min} .

Key words: Dirac monopole string, de Sitter space-time, Dirac equation, Wigner D -function, hypergeometric function.

Introduction

De Sitter and anti de Sitter geometrical models are given steady attention in the context of developing quantum theory in a curved space-time – for instance, see in [1, 2]. In particular, the problem of description of the particles with different spins on these curved backgrounds has a long history [3–36]. Here we will be interested mostly in treating the Dirac equation in de Sitter model.

In the present paper, the influence of the Dirac monopole string on the spin 1/2 particle in de Sitter cosmological model is investigated. Such a problem for spinless particle in the flat Minkowski space was first considered by Dirac [37] and Tamm [38]; Harish-Chandra [39] obtained the exact solution of Dirac equation for electron interacting with magnetic-monopole field. Instead of spinor monopole harmonics, the technique of Wigner D -functions is used. After separation of the variables radial equation have been solved exactly in terms of hypergeometric functions. The complete set of spinor wave solutions $\Psi_{\varepsilon,j,m,\lambda}(t, r, \theta, \phi)$ has been constructed. Special attention is given to

treating the states of minimal values for total angular momentum quantum number j_{\min} , these states turn to be much more complicated than in the flat Minkowski space.

1. Dirac particle in de Sitter space

The Dirac equation (the notation according to [40] is used)

$$\left[i\gamma^c (e_{(c)}^\alpha \partial_\alpha + \frac{1}{2} \sigma^{ab} \gamma_{abc}) - M \right] \Psi = 0 \quad (1)$$

in static coordinates and tetrad of the Sitter space (let $\Phi = 1 - r^2$)

$$dS^2 = \Phi dt^2 - \frac{dr^2}{\Phi} - r^2 (d\theta^2 + \sin^2 \theta d\phi^2),$$

takes the form

$$\left[i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \left(\gamma^3 \partial_r + \frac{\gamma^1 \sigma^{31} + \gamma^2 j^3}{r} + \frac{\Phi'}{2\Phi} \gamma^0 \sigma^{03} \right) + \frac{1}{r} \Sigma_{\theta,\phi} - M \right] \Psi(x) = 0, \quad (2)$$

where

$$\Sigma_{\theta,\phi} = i\gamma^1 \partial_\theta + \gamma^2 \frac{i\partial + i\sigma^{12} \cos \theta}{\sin \theta}.$$

With a simplifying substitution we get

$$\Psi(x) = r^{-1} \Phi^{-1/4} \psi(x),$$

$$\left(i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \gamma^3 \partial_r + \frac{1}{r} \Sigma_{\theta, \phi} - M \right) \psi = 0.$$

Below the spinor basis will be used.

2. Separation of the variables

Let us introduce a Dirac string potential in the de Sitter space-time model. It is convenient to start with the monopole Abelian potential in the Schwinger's form for the flat Minkowski space [41]

$$A^\alpha(x) = \left(0, g \frac{(\vec{r} \times \vec{n})(\vec{r}\vec{n})}{r(r^2 - (\vec{r}\vec{n})^2)} \right). \quad (3)$$

Specifying $\vec{n} = (0, 0, 1)$ and translating the $A_\alpha(x)$ to the spherical coordinates, we get $A_\phi = g \cos \theta$. This potential A_ϕ obeys Maxwell equations in de Sitter space as well. Correspondingly, the Dirac equation in presence of this field A_ϕ takes the form

$$\left(i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i \sqrt{\Phi} \gamma^3 \partial_r + \frac{1}{r} \Sigma_{\theta, \phi}^k - M \right) \psi = 0,$$

where (below the notation $eg/\hbar c = k$ will be used)

$$\Sigma_{\theta, \phi}^k = i\gamma^1 \partial_\theta + \gamma^2 \frac{i\partial_\phi + (i\sigma^{12} - k) \cos \theta}{\sin \theta}. \quad (4)$$

As readily verified, the wave operator in (4) commutes with the following three ones

$$J_1^k = l_1 + \frac{(i\sigma^{12} - k) \cos \phi}{\sin \theta},$$

$$J_2^k = l_2 + \frac{(i\sigma^{12} - k) \sin \phi}{\sin \theta}, J_3^k = l_3 \quad (5)$$

which in turn obey the $su(2)$ Lie algebra. Clearly, this monopole situation comes entirely under the Schrödinger [42], and Pauli [43] approach; detailed treatment of the method was given recently in [44]; similar technique when treating the problem of any spin particle in magnetic pole was used previously in [45], though with no connection with tetrad formalism.

Corresponding to diagonalization of the

\vec{J}_k^2 and J_3^k , the function ψ is to be taken as ($D_\sigma \equiv D_{-m, \sigma}^j(\phi, \theta, 0)$ stands for Wigner functions [46])

$$\psi_{\varepsilon jm}^k(t, r, \theta, \phi) = e^{-i\varepsilon t} \begin{pmatrix} f_1 D_{k-1/2} \\ f_2 D_{k+1/2} \\ f_3 D_{k-1/2} \\ f_4 D_{k+1/2} \end{pmatrix}. \quad (6)$$

Further, with the use of recursive relations [46] we find how the $\Sigma_{\theta, \phi}^k$ acts on $\psi_{\varepsilon jm}^k$

$$\Sigma_{\theta, \phi}^k \psi_{\varepsilon jm}^k = i \sqrt{(j+1/2)^2 - k^2} e^{-i\varepsilon t} \times \begin{pmatrix} -f_4 D_{k-1/2} (10) \\ +f_3 D_{k+1/2} (11) \\ +f_2 D_{k-1/2} (12) \\ -f_1 D_{k+1/2} \end{pmatrix}; \quad (7)$$

hereafter the factor $\sqrt{(j+1/2)^2 - k^2}$ will be referred to as ν . For the $f_i(r)$, the radial system derived is

$$\frac{\varepsilon}{\sqrt{\Phi}} f_3 - i \sqrt{\Phi} \frac{d}{dr} f_3 - i \frac{\nu}{r} f_4 - M f_1 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_4 + i \sqrt{\Phi} \frac{d}{dr} f_4 + i \frac{\nu}{r} f_3 - M f_2 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_1 + i \sqrt{\Phi} \frac{d}{dr} f_1 + i \frac{\nu}{r} f_2 - M f_3 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_2 - i \sqrt{\Phi} \frac{d}{dr} f_2 - i \frac{\nu}{r} f_1 - M f_4 = 0. \quad (8)$$

Else one operator can be diagonalized together with $i\partial_t, \vec{J}_k^2, J_3^k$: namely, a generalized Dirac operator

$$K^k = -i\gamma^0 \gamma^3 \Sigma_{\theta, \phi}^k. \quad (9)$$

From the eigenvalue equation $K^k \psi_{\varepsilon jm} = \lambda \psi_{\varepsilon jm}$ we can produce two possible values for this λ and the corresponding restrictions on $f_i(r)$

$$\lambda = -\delta \sqrt{(j+1/2)^2 - k^2},$$

$$f_4 = \delta f_1, f_3 = \delta f_2. \quad (10)$$

Correspondingly, the system (8) reduces to

$$\left(\sqrt{\Phi} \frac{d}{dr} + \frac{\nu}{r}\right) f + \left(\frac{\varepsilon}{\sqrt{\Phi}} + \delta M\right) g = 0,$$

$$\left(\sqrt{\Phi} \frac{d}{dr} - \frac{\nu}{r}\right) g - \left(\frac{\varepsilon}{\sqrt{\Phi}} - \delta M\right) f = 0, \quad (11)$$

we have translated equations to new functions

$$f = (f_1 + f_2)/\sqrt{2}, \quad g = (f_1 - f_2)i\sqrt{2}.$$

Note the quantization rules:

$$\frac{eg}{hc} = \pm 1/2, \pm 1, \pm 3/2, \dots;$$

$$j = |k| - 1/2, |k + 1/2|, |k| + 3/2, \dots \quad (12)$$

The case of minimal value $j_{\min} = |k| - 1/2$ must be separated and treated in a special way:

$$\psi_{j_{\min}}^{k>0}(x) = e^{-i\epsilon t} \begin{vmatrix} f_1(r)D_{k-1/2} \\ 0(20) \\ f_3(r)D_{k-1/2} \\ 0 \end{vmatrix};$$

$$\psi_{j_{\min}}^{k<0}(x) = e^{-i\epsilon t} \begin{vmatrix} 0 \\ f_2(r)D_{k+1/2} \\ 0 \\ f_4(r)D_{k+1/2} \end{vmatrix}, \quad (13)$$

and the relation $\Sigma_{\theta,\phi} \Psi_{j_{\min}} = 0$ holds. Thus, at every k , the j_{\min} -equation has the same form

$$\left(i \frac{\gamma^0}{\sqrt{\Phi}} \partial_t + i\gamma^3 \sqrt{\Phi} \left(\partial_r + \frac{1}{r}\right) - M\right) \psi_{j_{\min}} = 0;$$

which leads to the same radial system

$$k = +1/2, +1, \dots$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_3 - i\sqrt{\Phi} \frac{d}{dr} f_3 - Mf_1 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_1 + i\sqrt{\Phi} \frac{d}{dr} f_1 - Mf_3 = 0; \quad (14)$$

$$k = -1/2, -1, \dots$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_4 + i\sqrt{\Phi} \frac{d}{dr} f_4 - Mf_2 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_2 - i\sqrt{\Phi} \frac{d}{dr} f_2 - Mf_4 = 0. \quad (15)$$

3. Radial equations in the case j_{\min}

For brevity, let us examine only the case of the minimal value of j :

$$k = +1/2, +1, \dots$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_3 - i\sqrt{\Phi} \frac{d}{dr} f_3 - Mf_1 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_1 + i\sqrt{\Phi} \frac{d}{dr} f_1 - Mf_3 = 0; \quad (16)$$

from whence for new functions $h = (f_1 + f_3)/\sqrt{2}$, $g = (f_1 - f_3)/i\sqrt{2}$ we derive $k = +1/2, +1, \dots$

$$\sqrt{\Phi} \frac{d}{dr} h + \left(\frac{\varepsilon}{\sqrt{\Phi}} + M\right) g = 0,$$

$$\sqrt{\Phi} \frac{d}{dr} g - \left(\frac{\varepsilon}{\sqrt{\Phi}} - M\right) h = 0. \quad (17)$$

In the same manner for another case we have

$$k = -1/2, -1, \dots$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_4 + i\sqrt{\Phi} \frac{d}{dr} f_4 - Mf_2 = 0,$$

$$\frac{\varepsilon}{\sqrt{\Phi}} f_2 - i\sqrt{\Phi} \frac{d}{dr} f_2 - Mf_4 = 0; \quad (18)$$

where $g = (f_2 + f_4)/\sqrt{2}$, $h = (f_2 - f_4)/i\sqrt{2}$ we obtain

$$\sqrt{\Phi} \frac{d}{dr} h + \left(\frac{\varepsilon}{\sqrt{\Phi}} - M\right) g = 0,$$

$$\sqrt{\Phi} \frac{d}{dr} g - \left(\frac{\varepsilon}{\sqrt{\Phi}} + M\right) h = 0. \quad (19)$$

To exclude additional non-physical singular points, let us perform special transformation over the functions

$$g + h = e^{-i\rho^2} (F + G),$$

$$g - h = e^{+i\rho^2} (F - G). \quad (20)$$

After simple calculation we arrive at

$$k = +1/2, +3/2, \dots$$

$$\left(\frac{d}{d\rho} - i\varepsilon \frac{\sin \rho}{\cos \rho}\right) F +$$

$$+ \left(+\varepsilon + M - \frac{i}{2}\right) G = 0,$$

$$\left(\frac{d}{d\rho} + i\varepsilon \frac{\sin \rho}{\cos \rho}\right) G +$$

$$+ \left(-\varepsilon + M - \frac{i}{2}\right) F = 0; \quad (21)$$

$$\begin{aligned}
 k = -1/2, -3/2, \dots \\
 \left(\frac{d}{d\rho} - i\varepsilon \frac{\sin \rho}{\cos \rho} \right) G + \\
 + \left(+\varepsilon - M - \frac{i}{2} \right) H = 0, \\
 \left(\frac{d}{d\rho} + i\varepsilon \frac{\sin \rho}{\cos \rho} \right) H + \\
 + \left(-\varepsilon - M - \frac{i}{2} \right) G = 0. \quad (22)
 \end{aligned}$$

The difference between (21) and (22) consists in the only change $M \longleftrightarrow -M$.

In particular, the system (21) being translated to the variable z

$$\sin \rho = \sqrt{z}, \quad \frac{d}{d\rho} = 2\sqrt{z(1-z)} \frac{d}{dz},$$

will take the form

$$\begin{aligned}
 \sqrt{z(1-z)} \left(\frac{d}{dz} - \frac{i\varepsilon/2}{1-z} \right) F + \\
 + \frac{M + \varepsilon - i/2}{2} G = 0, \\
 \sqrt{z(1-z)} \left(\frac{d}{dz} + \frac{i\varepsilon/2}{1-z} \right) G + \\
 + \frac{M - \varepsilon - i/2}{2} F = 0. \quad (23)
 \end{aligned}$$

From (23) it follows 2-nd order differential equations for F and G respectively

$$\begin{aligned}
 z(1-z) \frac{d^2 F}{dz^2} + \left(\frac{1}{2} - z \right) \frac{dF}{dz} + \\
 + \left[-\frac{1}{4} \left(M - \frac{i}{2} \right)^2 + \frac{\varepsilon(\varepsilon - i)}{4(1-z)} \right] F = 0, \\
 z(1-z) \frac{d^2 G}{dz^2} + \left(\frac{1}{2} - z \right) \frac{dG}{dz} + \\
 + \left[-\frac{1}{4} \left(M - \frac{i}{2} \right)^2 + \frac{\varepsilon(\varepsilon + i)}{4(1-z)} \right] G = 0.
 \end{aligned}$$

Let us introduce substitutions

$$F = z^A (1-z)^B \bar{F}(z), \quad G = z^K (1-z)^L \bar{G}(z).$$

At A and B taken accordingly

$$A = \frac{1}{2}, 0, \quad B = -\frac{i\varepsilon}{2}, \frac{1+i\varepsilon}{2} \quad (24)$$

equation for \bar{F} looks

$$\begin{aligned}
 z(1-z) \frac{d^2 \bar{F}}{dz^2} + \\
 + \left[2A + \frac{1}{2} - (2A + 2B + 1)z \right] \frac{d\bar{F}}{dz} + \\
 + \left[-\frac{1}{4} \left(M - \frac{i}{2} \right)^2 - (A+B)^2 \right] \bar{F} = 0
 \end{aligned}$$

which is of hypergeometric type with parameters

$$a = A + B + \frac{iM + 1/2}{2},$$

$$b = A + B - \frac{iM + 1/2}{2}, \quad c = 2A + 1/2.$$

To have solutions non-vanishing in the origin $z=0$, we take zero $A=0$. Thus there arise two sorts of solutions depending on a chosen B (in each case two linearly independent solutions, regular and singular in the origin, are written down)

the first

$$\begin{aligned}
 c = +1/2, \quad \bar{F}_{non-zero}^{(1)}(z) = F(a, b, c; z), \\
 \bar{F}_{zero}^{(1)} = z^{1-c} F(a+1-c, b+1-c, 2-c; z), \\
 a = \frac{-i\varepsilon}{2} + \frac{iM + 1/2}{2}, \quad b = \frac{-i\varepsilon}{2} - \frac{iM + 1/2}{2};
 \end{aligned}$$

the second

$$\begin{aligned}
 \gamma = +1/2, \quad \bar{F}_{non-zero}^{(2)} = F(\alpha, \beta, \gamma; z), \quad \bar{F}_{zero}^{(2)} = \\
 = z^{1-\gamma} F(\alpha+1-\gamma, \beta+1-\gamma, 2-\gamma; z),
 \end{aligned}$$

where

$$\begin{aligned}
 \alpha = \frac{1+i\varepsilon}{2} + \frac{iM + 1/2}{2}, \\
 \beta = \frac{1+i\varepsilon}{2} - \frac{iM + 1/2}{2}.
 \end{aligned}$$

Now let us turn to equation for \bar{G} ; at K and L chosen according to

$$K = \frac{1}{2}, 0, \quad L = \frac{i\varepsilon}{2}, \frac{1-i\varepsilon}{2} \quad (25)$$

it looks

$$\begin{aligned}
 z(1-z) \frac{d^2 \bar{G}}{dz^2} + \\
 + \left[2K + \frac{1}{2} - (2K + 2L + 1)z \right] \frac{d\bar{G}}{dz} + \\
 + \left[-\frac{1}{4} \left(M - \frac{i}{2} \right)^2 - (K+L)^2 \right] \bar{G} = 0, \quad (36)
 \end{aligned}$$

which is of hypergeometric type

$$a' = K + L + \frac{iM + 1/2}{2},$$

$$b' = K + L - \frac{iM + 1/2}{2}, \quad c' = 2K + \frac{1}{2}.$$

We start with solutions non-vanishing in the origin $z=0$, we take zero $K=0$. Thus there arise two sorts of solutions depending on a chosen B (in each case two linearly independent solutions, regular and singular in the origin, are written down)

the first

$$c' = +1/2, \quad \bar{G}_{non-zero}^{(1)} = F(a', b', c'; z), \quad \bar{G}_{zero}^{(1)} = z^{1-c'} F(a'+1-c', b'+1-c', 2-c'; z),$$

$$a' = \frac{+i\varepsilon}{2} + \frac{iM + 1/2}{2}, \quad b' = \frac{+i\varepsilon}{2} - \frac{iM + 1/2}{2};$$

the second

$$\gamma' = +1/2, \quad \bar{G}_{non-zero}^{(2)}(z) = F(\alpha', \beta', \gamma'; z),$$

$$\bar{G}_{zero}^{(2)} = z^{1-\gamma'} F(\alpha'+1-\gamma', \beta'+1-\gamma', 2-\gamma'; z), \quad \alpha' = \frac{1-i\varepsilon}{2} + \frac{iM + 1/2}{2},$$

$$\beta' = \frac{1-i\varepsilon}{2} - \frac{iM + 1/2}{2}.$$

Thus, we have constructed the following four regular solutions

$$F_{non-zero}^{(1)}, G_{non-zero}^{(1)}, G_{non-zero}^{(2)}.$$

Due to the known identity for hypergeometric functions

$$F(A, B, C; z) = (1-z)^{C-A-B} F(C-A, C-B, C; z)$$

we readily conclude that there exist only two different ones

$$F_{non-zero}^{(1)} = F_{non-zero}^{(2)},$$

$$G_{non-zero}^{(1)} = G_{non-zero}^{(2)}. \quad (26)$$

The same is true for zero-solutions

$$F_{zero}^{(1)} = F_{zero}^{(2)}, \quad G_{zero}^{(1)} = G_{zero}^{(2)}. \quad (27)$$

Assuming relationship

$$2\sqrt{z(1-z)} \left(\frac{d}{dz} - \frac{i\varepsilon/2}{1-z} \right) F_{non-zero}^{(1)} + (M + \varepsilon - i/2) G_{zero}^{(2)} = 0.$$

we readily derive

$$aF_0^{non-zero} + icG_0^{zero} = 0. \quad (28)$$

And similarly, we can expect

$$2\sqrt{z(1-z)} \left(\frac{d}{dz} + \frac{i\varepsilon/2}{1-z} \right) G_{non-zero}^{(1)} + (M - \varepsilon - i/2) F_{zero}^{(2)} = 0,$$

so that

$$a'G_0^{non-zero} + ic'F_0^{zero} = 0. \quad (29)$$

4. Standing and running waves, $j = j_{\min}$

Let write down results we need to proceed further

$$F_{non-zero} = (1-z)^{-i\varepsilon/2} U_1,$$

$$F_{zero} = (1-z)^{-i\varepsilon/2} U_5,$$

$$G_{non-zero} = (1-z)^{+i\varepsilon/2} V_1,$$

$$G_{zero} = (1-z)^{+i\varepsilon/2} V_5, \quad (30)$$

$$U_1 = F(a, b, c, z), \quad V_1 = F(a', b', c', z),$$

$$a = \frac{-i\varepsilon}{2} + \frac{iM + 1/2}{2},$$

$$b = \frac{-i\varepsilon}{2} - \frac{iM + 1/2}{2}, \quad c = 1/2,$$

$$a' = \frac{+i\varepsilon}{2} + \frac{iM + 1/2}{2},$$

$$b' = \frac{+i\varepsilon}{2} - \frac{iM + 1/2}{2}, \quad c' = 1/2.$$

We readily derive

$$F_{non-zero} = (1-z)^{-i\varepsilon/2} U_1$$

$$= \frac{\Gamma(c)\Gamma(c-a-b)}{\Gamma(c-a)\Gamma(c-b)} F_{out} + \frac{\Gamma(c)\Gamma(a+b-c)}{\Gamma(a)\Gamma(b)} F_{in},$$

$$F_{zero} = (1-z)^{-i\varepsilon/2} U_5$$

$$= \frac{\Gamma(2-c)\Gamma(c-a-b)}{\Gamma(1-a)\Gamma(1-b)} F_{out} + \frac{\Gamma(2-c)\Gamma(a+b-c)}{\Gamma(a+1-c)\Gamma(b+1-c)} F_{in},$$

where

$$F_{out} = (1-z)^{-i\epsilon/2} U_2, F_{in} = (1-z)^{-i\epsilon/2} U_6, \quad h(r) = He^{\gamma r}, \quad g(r) = Ge^{\gamma r} \quad (35)$$

$$G_{non-zero} = (1-z)^{+i\epsilon/2} V_1 \\ = \frac{\Gamma(c)\Gamma(c-a-b)}{\Gamma(c-a)\Gamma(c-b)} G_{in} + \\ + \frac{\Gamma(c)\Gamma(a+b-c)}{\Gamma(a)\Gamma(b)} G_{out},$$

$$G_{zero} = (1-z)^{+i\epsilon/2} V_5 \\ = \frac{\Gamma(2-c')\Gamma(c'-a'-b')}{\Gamma(1-a')\Gamma(1-b')} G_{in} + \\ + \frac{\Gamma(2-c')\Gamma(a'+b'-c')}{\Gamma(a'+1-c')\Gamma(b'+1-c')} G_{out},$$

where

$$G_{in} = (1-z)^{+i\epsilon/2} V_2, G_{out} = (1-z)^{+i\epsilon/2} V_6.$$

5. Discussion and conclusions

To understand better the situation, let us consider the case of minimal j_{min} in the limit of vanishing curvature. It is convenient to start with the first order systems for minimal values j_{min} in the case of Minkowski space:

$$k = +1/2, +1, \dots$$

$$\epsilon f_3 - i \frac{d}{dr} f_3 - M f_1 = 0, \\ \epsilon f_1 + i \frac{d}{dr} f_1 - M f_3 = 0; \quad (31)$$

$$k = -1/2, -1, \dots$$

$$\epsilon f_4 + i \frac{d}{dr} f_4 - M f_2 = 0, \\ \epsilon f_2 - i \frac{d}{dr} f_2 - M f_4 = 0. \quad (32)$$

Let us detail the case of positive $k = +1/2, +1, \dots$. With notation

$$\frac{f_1 + f_3}{\sqrt{2}} = h(r), \frac{f_1 - f_3}{i\sqrt{2}} = g(r); \quad (33)$$

relevant equations are

$$\frac{d}{dr} h + (\epsilon + M)g = 0, \\ \frac{d}{dr} g - (\epsilon - M)h = 0. \quad (34)$$

Further, with the substitutions

we get (first let it be $(\epsilon^2 - M^2) > 0$)

$$\gamma^2 = -(\epsilon^2 - M^2) = -p^2, \\ \gamma = +ip, -ip, G\gamma - (\epsilon - M)H = 0. \quad (36)$$

Thus we have two linearly independent solutions

$$h_1(r) = H_1 e^{+ipr}, g_1(r) = G_1 e^{+ipr}, \\ G_1 = \frac{\epsilon - M}{ip} H_1 \quad (37)$$

and

$$h_2(r) = H_2 e^{-ipr}, g_2(r) = G_2 e^{-ipr}, \\ G_2 = \frac{\epsilon - M}{-ip} H_2; \quad (38)$$

for simplicity, we will take $H_1 = H_2 = 1$. It is convenient to introduce linear combinations of these solutions

$$\frac{h_1(r) + h_2(r)}{2} = \cos pr, \\ \frac{g_1(r) + g_2(r)}{2} = \frac{\epsilon - M}{p} \sin pr; \quad (39)$$

$$\frac{h_1(r) - h_2(r)}{2i} = \sin pr, \\ \frac{g_1(r) - g_2(r)}{2i} = \frac{\epsilon - M}{-p} \cos pr. \quad (40)$$

Now let us specify the case $(\epsilon^2 - M^2) < 0$:

$$\gamma^2 = -(\epsilon^2 - M^2) \equiv +q^2, \\ \gamma = +q, -q, \\ G\gamma - (\epsilon - M)H = 0. \quad (41)$$

We have two linearly independent solutions

$$h_1(r) = H_1 e^{+qr}, g_1(r) = G_1 e^{+qr}, \\ G_1 = \frac{\epsilon - M}{q} H_1; \quad (42) \\ h_2(r) = H_2 e^{-qr}, g_2(r) = G_2 e^{-qr},$$

$$G_2 = \frac{\varepsilon - M}{-q} H_2. \quad (43)$$

Below $H_1 = H_2 = 1$. We can introduce two linear combinations of these solutions

$$\frac{h_1(r) + h_2(r)}{2} = \cosh qr, \quad (44)$$

$$\frac{g_1(r) + g_2(r)}{2} = \frac{\varepsilon - M}{q} \sinh qr,$$

$$\frac{h_1(r) - h_2(r)}{2} = \sinh qr,$$

$$\frac{g_1(r) - g_2(r)}{2} = \frac{\varepsilon - M}{q} \cosh qr. \quad (45)$$

Evidently, above constructed solutions in de Sitter model provide us with generalizations of these in Minkowski space. It may be verified additionally by direct limiting process when $\rho \rightarrow \infty$. To this end, let us translate solutions in de Sitter space to usual units (ρ is the curvature radius, E is the energy, c is the light velocity, m is the electron mass)

$$F_{non-zero} = \left(1 - \frac{R^2}{\rho^2}\right)^{-i\frac{E\rho}{2\hbar c}} F\left(a, b, c; \frac{R^2}{\rho^2}\right),$$

$$F_{zero} = R \left(1 - \frac{R^2}{\rho^2}\right)^{+i\frac{E\rho}{2\hbar c} + 1/2} \times$$

$$\times F\left(a+1-c, b+1-c, 2-c; \frac{R^2}{\rho^2}\right),$$

$$G_{non-zero} = \left(1 - \frac{R^2}{\rho^2}\right)^{+i\frac{E\rho}{2\hbar c}} F\left(a', b', c; \frac{R^2}{\rho^2}\right),$$

$$G_{zero} = R \left(1 - \frac{R^2}{\rho^2}\right)^{-i\frac{E\rho}{2\hbar c} + 1/2} \times$$

$$\times F\left(a'+1-c, b'+1-c, 2-c; \frac{R^2}{\rho^2}\right),$$

Parameters of hypergeometric functions are given by

$$c = \frac{1}{2}, a = \frac{1}{2} \left[+1/2 + i \left(\frac{mc\rho}{\hbar} - \frac{E\rho}{c\hbar} \right) \right],$$

$$b = \frac{1}{2} \left[-i \left(\frac{mc\rho}{\hbar} + \frac{E\rho}{c\hbar} \right) - 1/2 \right],$$

$$c' = \frac{1}{2}, a' = \frac{1}{2} \left[+1/2 + i \left(\frac{mc\rho}{\hbar} + \frac{E\rho}{c\hbar} \right) \right],$$

$$b' = \frac{1}{2} \left[-i \left(\frac{mc\rho}{\hbar} - \frac{E\rho}{c\hbar} \right) - 1/2 \right].$$

Let us examine the limiting procedure at $\rho \rightarrow \infty$ in $F(a, b, c; R^2/\rho^2)$. Because

$$\frac{1}{1!} \frac{ab R^2}{c \rho^2} \rightarrow -\frac{1}{2!} (pR)^2,$$

$$\frac{1}{2!} \frac{a(a+1)b(b+1) R^2}{c(c+1) \rho^2} \rightarrow +\frac{(pR)^4}{4!},$$

$$\frac{1}{3!} \frac{a(a+1)(a+2)b(b+1)(b+2) R^2}{c(c+1)(c+2) \rho^2}$$

$$\rightarrow -\frac{(pR)^6}{6!},$$

and so on, we obtain the limiting relations

$$\lim_{\rho \rightarrow \infty} F\left(a, b, c; \frac{R^2}{\rho^2}\right) = \cos pr \quad \Rightarrow$$

$$\lim_{\rho \rightarrow \infty} F_{non-zero}(R) = \cos pr,$$

$$\lim_{\rho \rightarrow \infty} G_{non-zero}(R) = \cos pr. \quad (46)$$

In the same manner, let us examine the function

$$RF\left(a+1-c, b+1-c, 2-c; \frac{R^2}{\rho^2}\right),$$

$$A = a+1-c = \frac{3/2 + i(M + \varepsilon)}{2},$$

$$B = b+1-c = \frac{1/2 - i(M - \varepsilon)}{2}, C = 3/2.$$

Taking into account

$$\frac{AB}{C} \rightarrow -\frac{1}{3!} (pR)^2,$$

$$\frac{1}{2!} \frac{A(A+1)B(B+1)}{C(C+1)} \rightarrow +\frac{1}{5!} (pR)^4,$$

$$\frac{1}{3!} \frac{A(A+1)(A+2)B(B+1)(B+2)}{C(C+1)(C+2)}$$

$$\rightarrow -\frac{1}{7!} (pR)^6,$$

and so on, we arrive at the relationships

$$\lim_{\rho \rightarrow \infty} pRF\left(a+1-c, b+1-c, 2-c; \frac{R^2}{\rho^2}\right) = \sin pR,$$

$$\lim_{\rho \rightarrow \infty} pRF_{zero} = \sin pR, \lim_{\rho \rightarrow \infty} pRG_{zero} = \sin pR.$$

Thus, indeed, solutions in de Sitter model are extensions of more simple and well-known ones in Minkowski space.

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ЧАСТИНКИ ЗІ СПІНОМ 1/2 У ПРИСУТНОСТІ АБЕЛЕВА МОНОПОЛЯ НА ФОНІ ПРОСТОРУ-ЧАСУ ДЕ СІТТЕРА

У космологічній моделі де Сіттера визначено потенціал діраківського монополя. Досліджено рівняння Дірака в присутності цього поля в статичних координатах простору де Сіттера. Замість спірних монопольних гармонік використано техніку D -функцій Вігнера. Після розділення змінних проведено детальний аналіз отриманих радіальних рівнянь, в термінах гіпергеометричних функцій побудовано чотири типи розв'язків: сингулярне, регулярне, біжучі хвилі, які сходяться і розбігаються. Знайдено повний набір хвильових розв'язків $\Psi_{\varepsilon, j, m, \lambda}(t, r, \theta, \phi)$, особливу увагу приділено дослідженню станів з мінімальним значенням повного кутового моменту j_{min} .

Ключові слова: потенціал діраківського монополя, простір-час де Сіттера, рівняння Дірака, D -функція Вігнера, гіпергеометричні функції.

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ЧАСТИЦЫ СО СПИНОМ $1/2$ В ПРИСУТСТВИИ АБЕЛЕВА МОНОПОЛЯ НА ФОНЕ ПРОСТРАНСТВА-ВРЕМЕНИ ДЕ СИТТЕРА

В космологической модели де Ситтера определен потенциал дираковского монополя. Уравнение Дирака в присутствии этого поля исследуется в статических координатах пространства де Ситтера. Вместо спинорных монополярных гармоник используется техника D -функций Вигнера. После разделения переменных проведен детальный анализ полученных радиальных уравнений, четыре типа решений: сингулярное, регулярное, расходящиеся и сходящиеся бегущие волны построены в терминах гипергеометрических функций. Найден полный набор волновых решений $\Psi_{\varepsilon, j, m, \lambda}(t, r, \theta, \phi)$, особое внимание уделено исследованию состояний с минимальным значением полного углового момента j_{min} .

Ключевые слова: потенциал дираковского монополя, пространство-время де Ситтера, уравнение Дирака, D -функция Вигнера, гипергеометрическая функция.